

SLR2000: current status & issues being worked
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Khaz ranging session
ILRS Workshop at Eastbourne
October 2005

Progress since Workshop in June 2004 (glacial – but in a positive direction)

- Point-ahead of transmit using Risley Prisms: checked out and working.
- Transceiver transformations documented (McGarry & Degnan, independently), verified, and parameters determined.
- New SigmaSpace Beam Expander and Iris are in the system, aligned and in the software. Both still need final checkout in the system.
- New photon-counting optical alignment technique using stars appears to do a better job of distributing light over the quadrants uniformly. Very important for closed-loop tracking.
- Replaced original detector with spare – appeared to have degraded.
- Started automated closed-loop tracking – some success – some problems.

*Work on SLR2000 was extremely slow March – September 2005
due to work on Transponder experiments*

Currently ongoing SLR2000 work

- SigmaSpace is working on design to optically shutter the detector.
- We are working on technique to change laser PRF to prevent collisions between outgoing & incoming pulses. Currently we are just blanking electronically.
- We plan to purchase a higher QE 4QMCP detector from Hamamatsu.
- We have analyzed the closed-loop tracking data and need to make some adjustments in the system.

Future work

- SLR2000 has been selected as primary ground station for LRO laser uplink. Will require the purchase of a higher energy 28Hz laser.

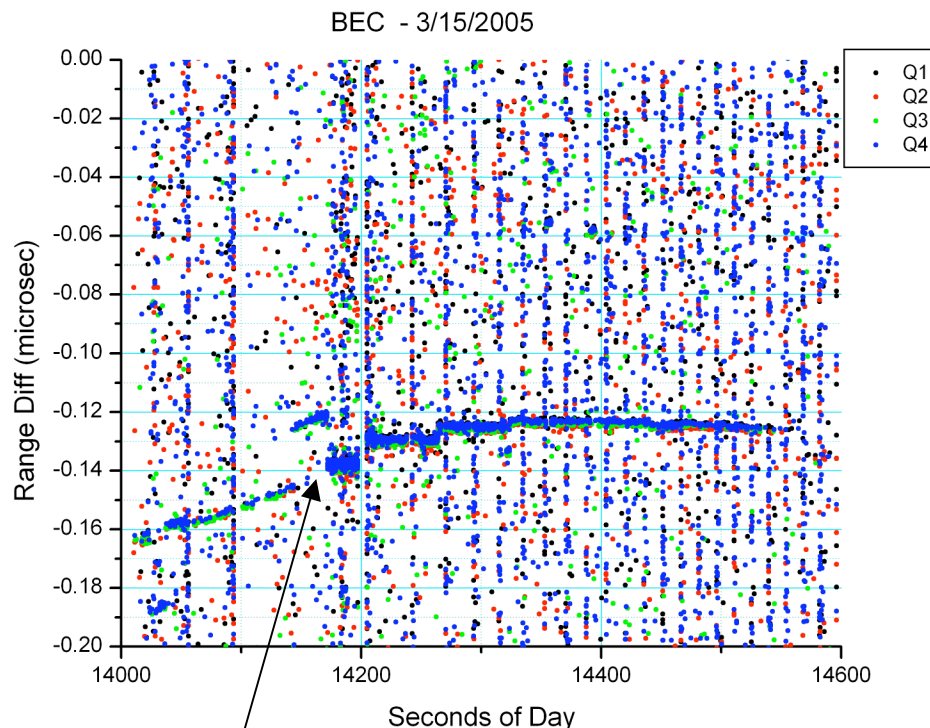
Issues related to khz ranging

Closed loop tracking: sometime it works and sometimes it doesn't!

- It is very important to get a uniform distribution of background noise across the four detectors and it is not easy. Probably need to use a software calibration of the quadrant levels to provide final correction.
- Signal must be centered on quadrants when telescope is directly pointed at target. Degnan conjectures that point-ahead may be causing backscattered light to favor one quadrant over others. This does appear to be happening in some of the passes, but not in others. ???
- We have a serious backscatter problem in SLR2000 which is causing problems in a lot of areas, but especially in degradation of the detector and most likely also in the closed-loop tracking. We must drastically reduce the backscatter effects through optical blanking and by changing the laser PRF.

Still not able to range to LAGEOS

- Returns from all satellites are generally weaker than theoretical.
- Until new beam expander, system wasn't properly focused, beam divergence was unknown and not easily changed. This should improve performance.



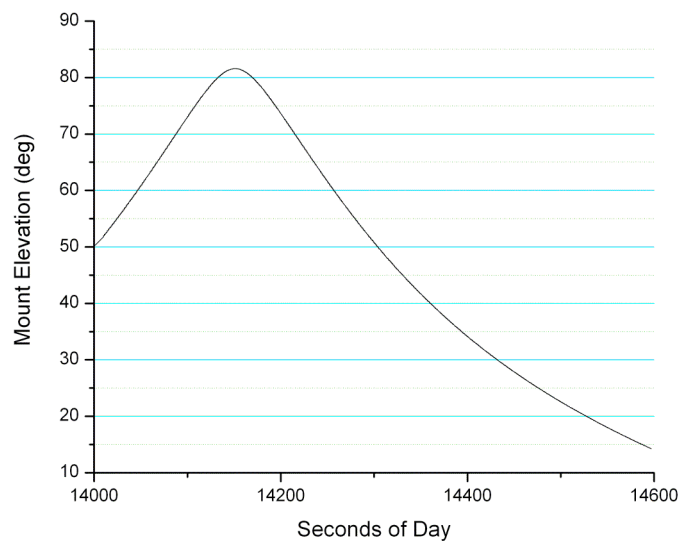
First closed-loop correction here

BEC pass taken on 3/14/2005 (GMT 3/15/2005 03:47).

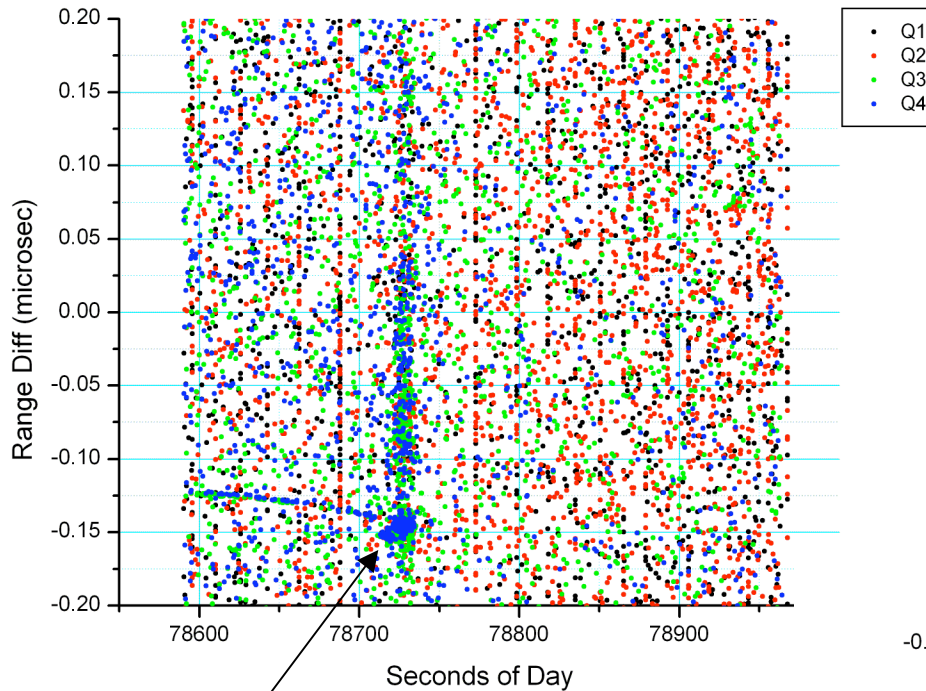
Problem with s/w recording of range-bias cause jumps in OMC plot.

Blanking causes periodic no-noise segments.

This pass was run with closed-loop activated and working. Point-ahead version of software.



STELLA - 3/15/2005



First closed-loop correction here

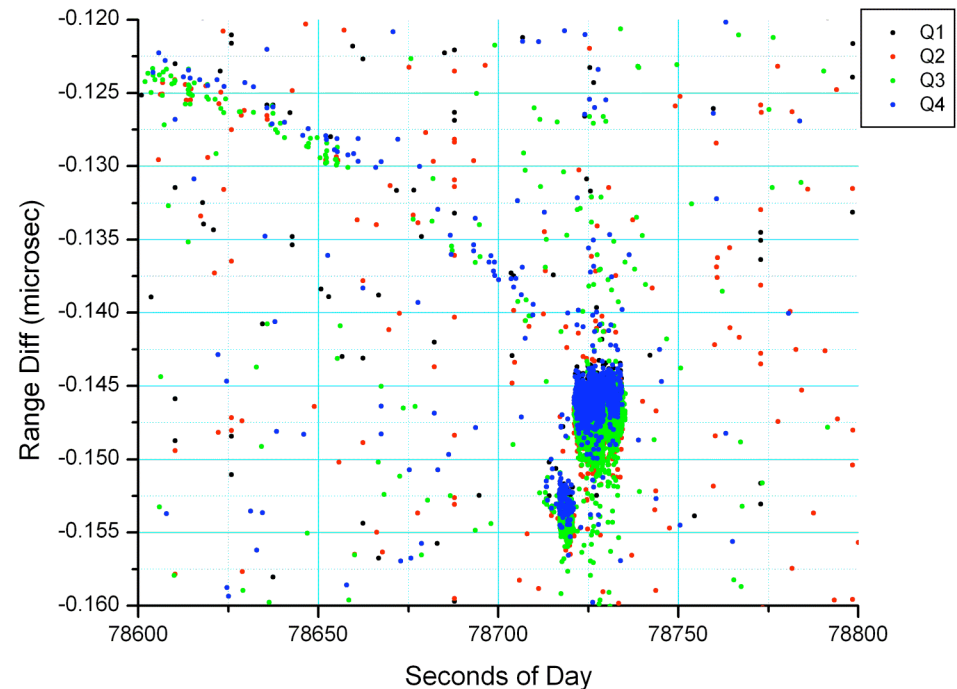
STELLA pass taken on 3/15/2005 (GMT 3/15/2005 21:46).

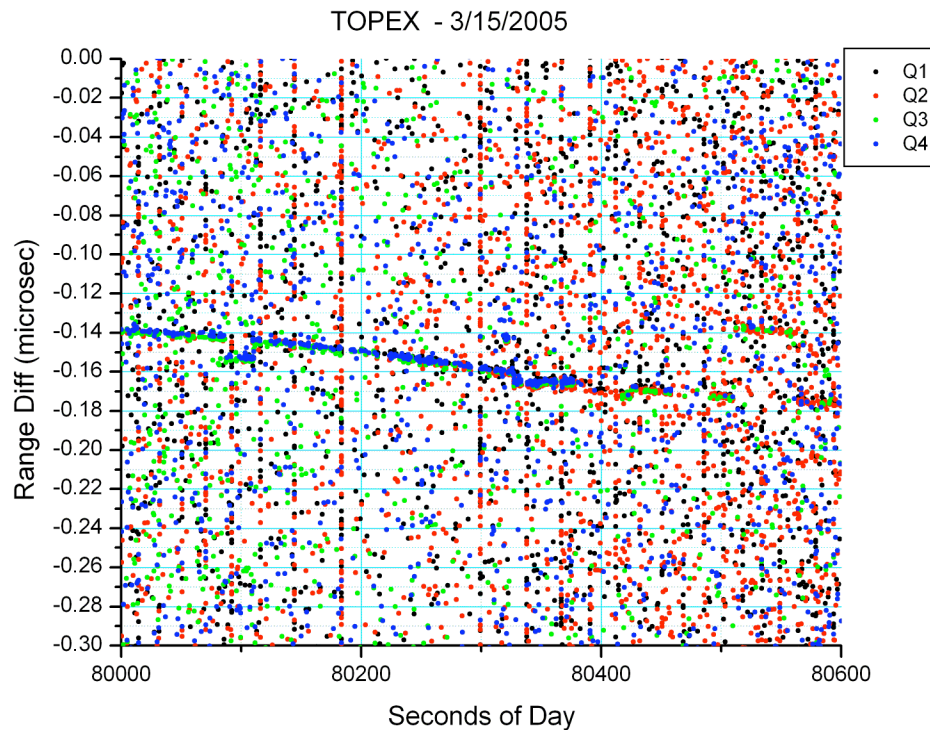
Problem with s/w recording of range-bias cause jumps in OMC plot.

Blanking still there – but daylight fills in gaps.

This pass was run with closed-loop activated. Point-ahead version of software.

STELLA - 3/15/2005



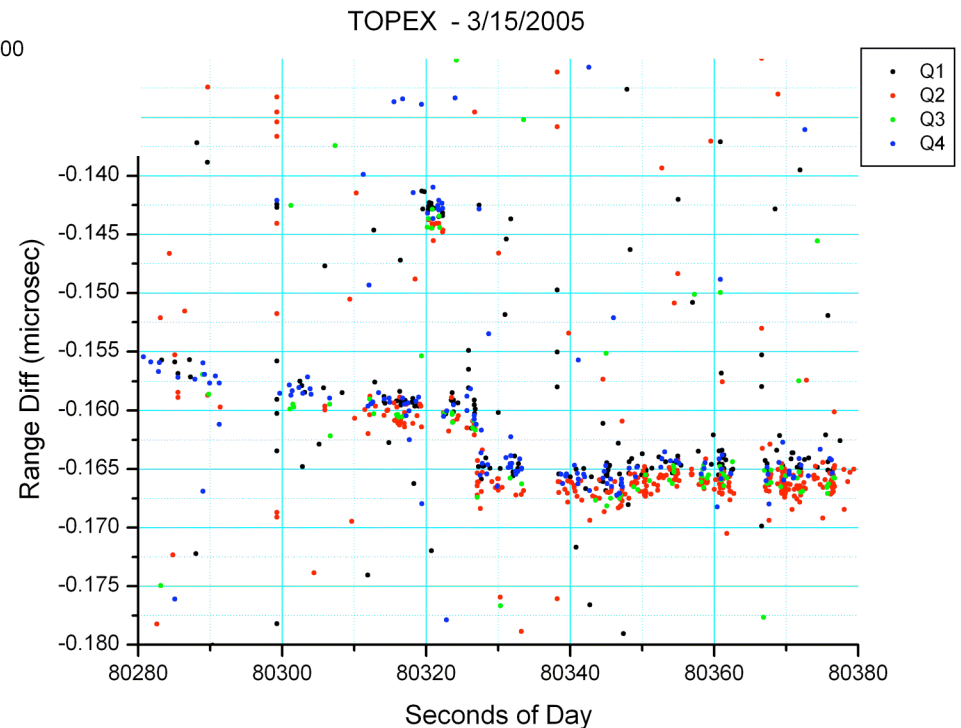


TOPEX pass taken on 3/15/2005 (GMT 3/15/2005 22:09).

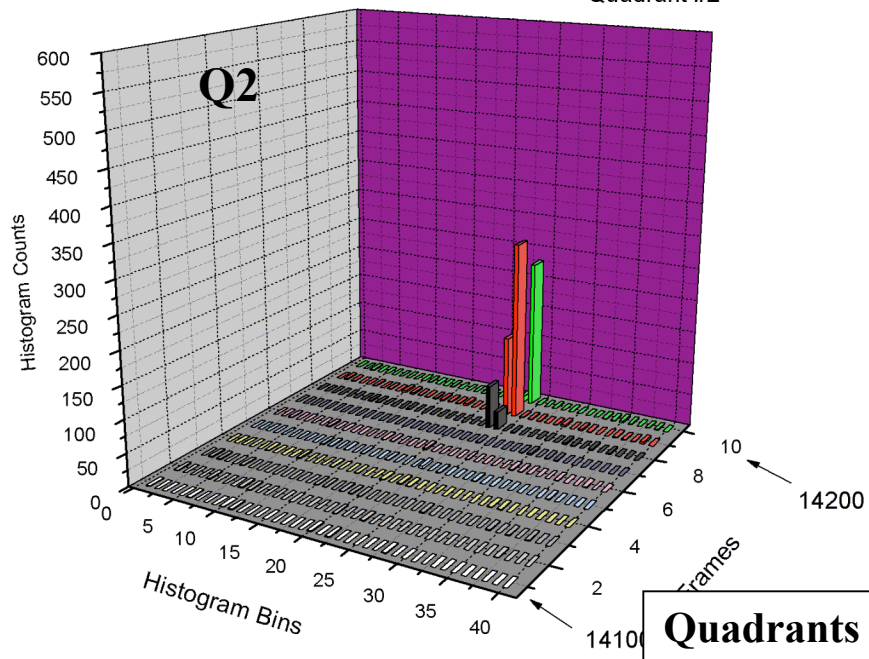
Problem with s/w recording of range-bias cause jumps in OMC plot.

Blanking causes periodic no-noise segments.

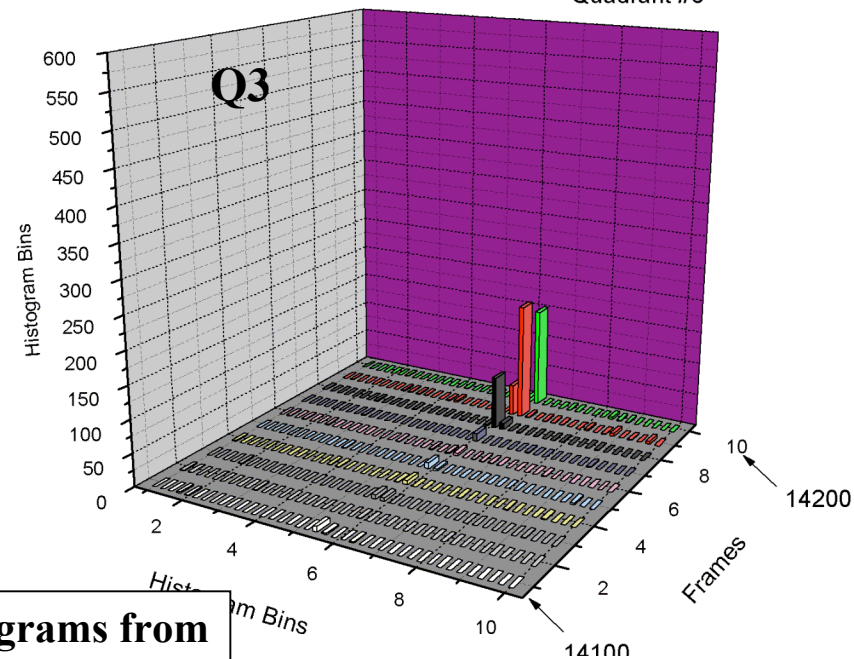
This pass was run with closed-loop activated. Point-ahead version of software.



Quadrant #2

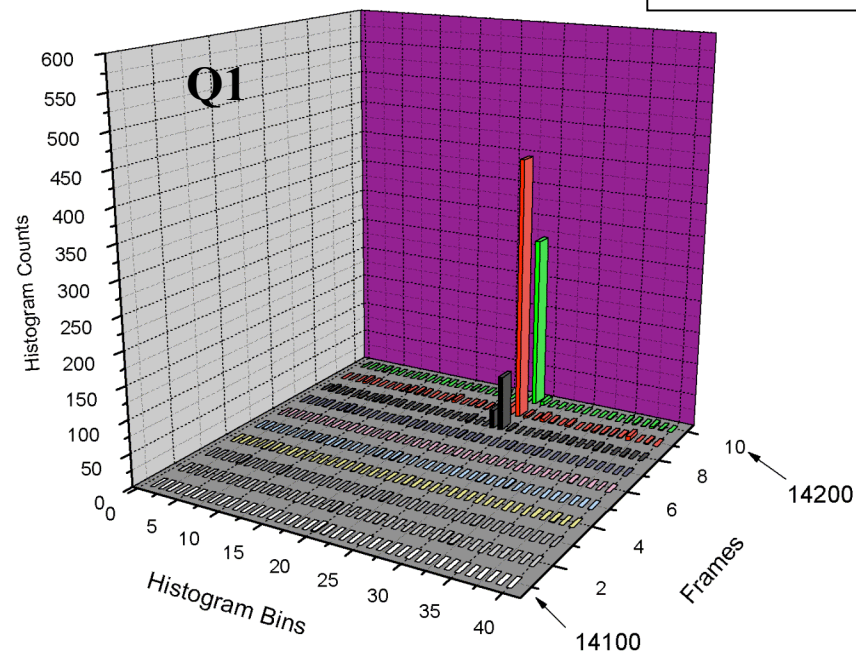
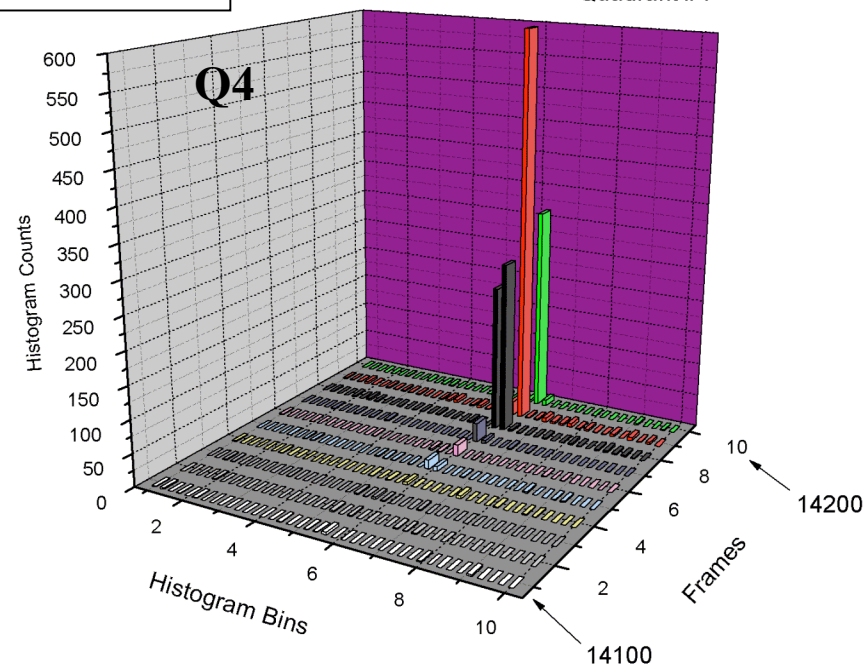


Quadrant #3



**Quadrants histograms from
BEC PASS versus time**

Quadrant #4



Potential new S2K detector:

Four quadrant MCP.

QE \geq 30%.

Price: \sim \$35k.

TECHNICAL INFORMATION

SPECIFICATION SHEET

Microchannel Plate Photomultiplier Tube

R4110U-74-M004B

GENERAL

Parameter	Ratings	Units
Spectral Response	280 to 720	nm
Wavelength of Maximum Response	480 to 550	nm
Input Window	Material	Borosilicate Glass
	Thickness	4.77
Photocathode	Material	GaAsP
	Minimum Effective Diameter	25
Microchannel Plate (MCP) Construction		2 Stages (Filmless)
Anode	Pattern	Matrix 4 ch
	Size (each ch)	8.5 x 8.5
Mechanical Size (H x W x L)		69 mm x 63 mm x 34.4 mm

RATINGS

Parameters	Ratings	Maximum Ratings	Units
Input Gate Pulse (Photocathode to MCPin)	Voltage	-200	V
	Width	100	ns
Supply Voltage (MCPin to Anode)	2200	-	V
Photocathode Current	-	10	nA
Average Current (Each Anode)	-	200	nA
Ambient Temperature	-	-50 to +50	°C

CHARACTERISTICS

Parameters	Min.	Typ.	Max.	Units
Photocathode Sensitivity	Luminous Sensitivity	400	700	-
	Quantum Efficiency at Peak	30	40	%
DC Current Amplification at -2200 V	1×10^3	3×10^3	-	-
Anode	Dark Counts at -2200 V (25 °C) (summed all anodes)	-	30000	s ⁻¹
Voltage Divider Current at -2200 V		-	100	uA
Photocathode Uniformity		50	-	%
MCP Gain Uniformity		50	-	%
Time Response (Gate on Operating)	Rise Time	-	200	ps
	Fall Time	-	300	
	T.T.S. (FWHM)	-	100	
Cross - Talk		-	8	%

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HAMAMATSU

Electron Tube Division
Department #3 Aug. 2005

